

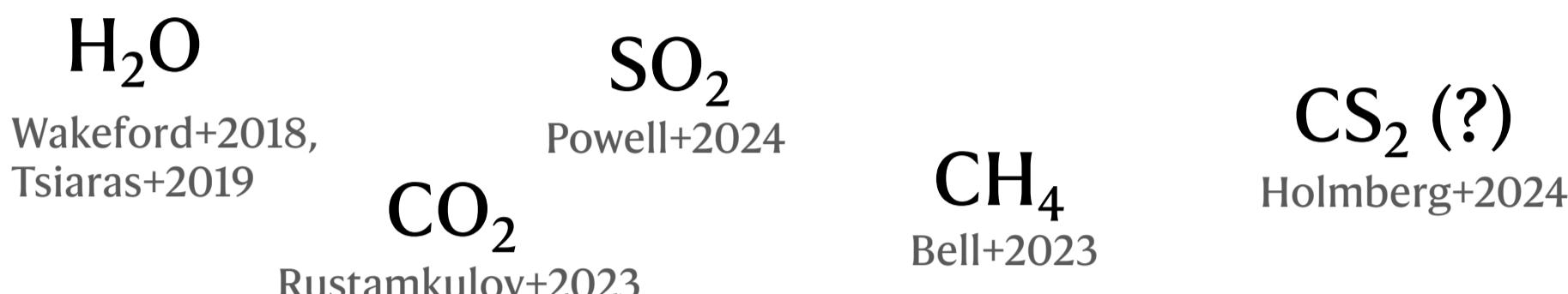


Chemical controls on the atmospheric stability of rocky exoplanets

(and why the C/O ratio is not enough)

Why chemistry?

A big variety of chemical species has been detected in exoplanet atmospheres



But the focus is very often on the C/O ratio or metallicity

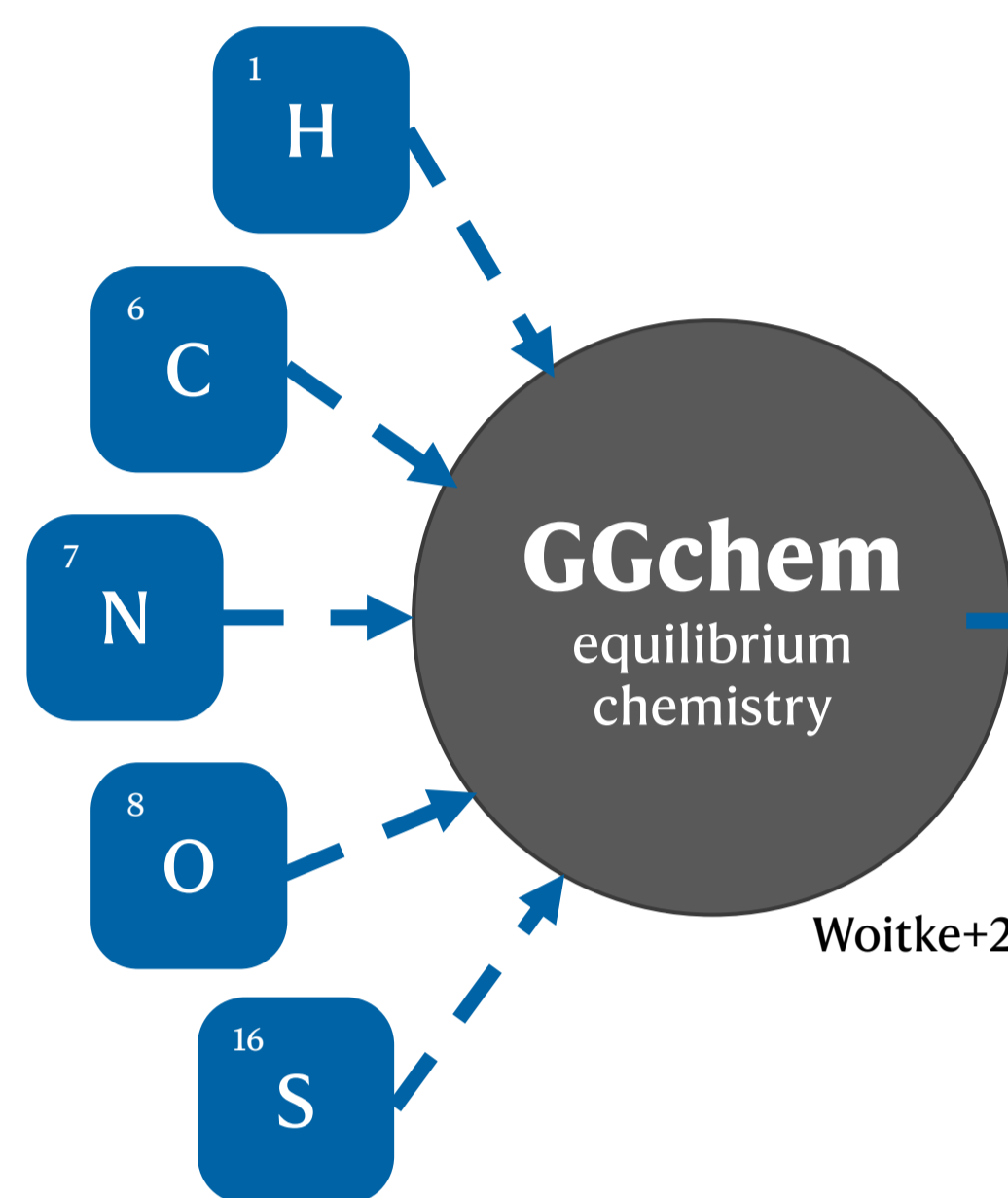
What about us?!

We want to see which chemical compositions are the most stable, and which are more easily lost



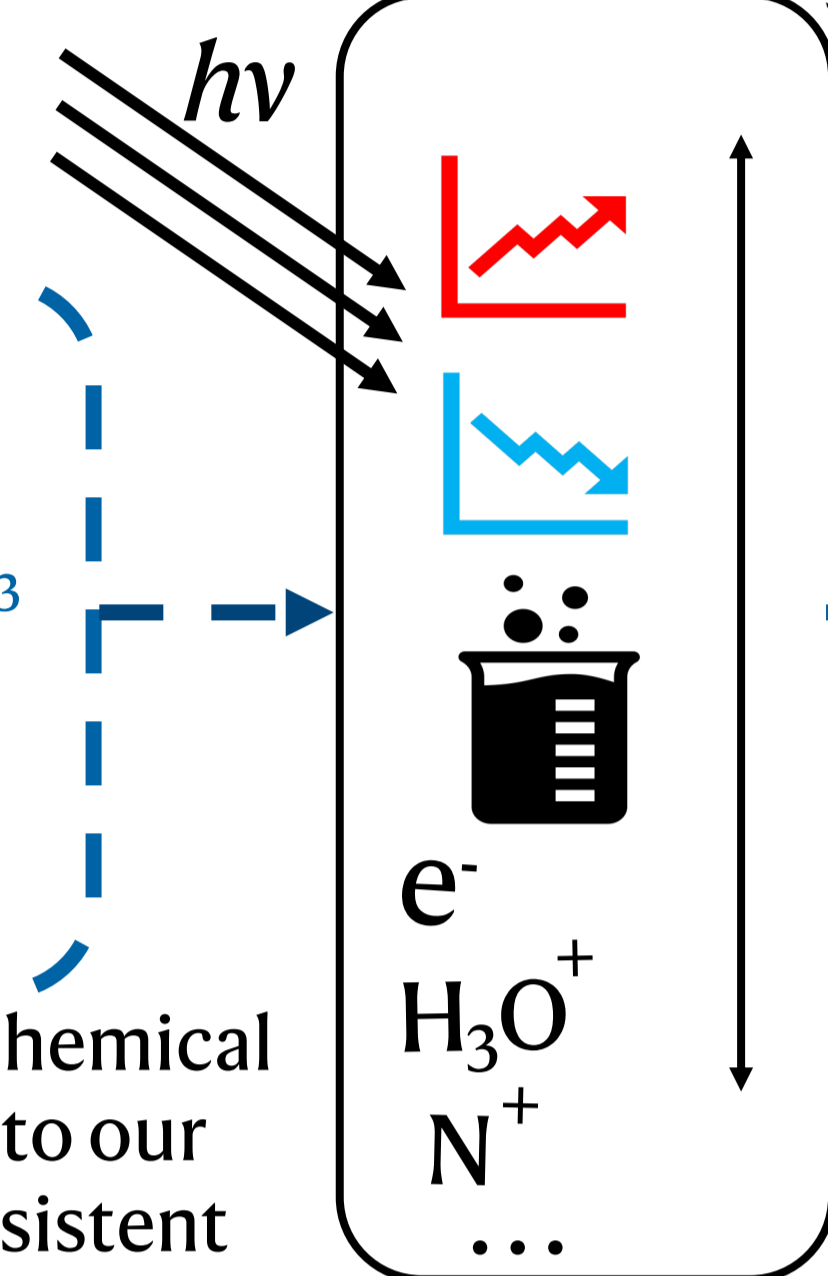
What did we model and how?

We vary these five input elemental fractions



Molecules from chemical equilibrium go into our in-house self-consistent atmospheric model

Kompot (Johnstone+2018)



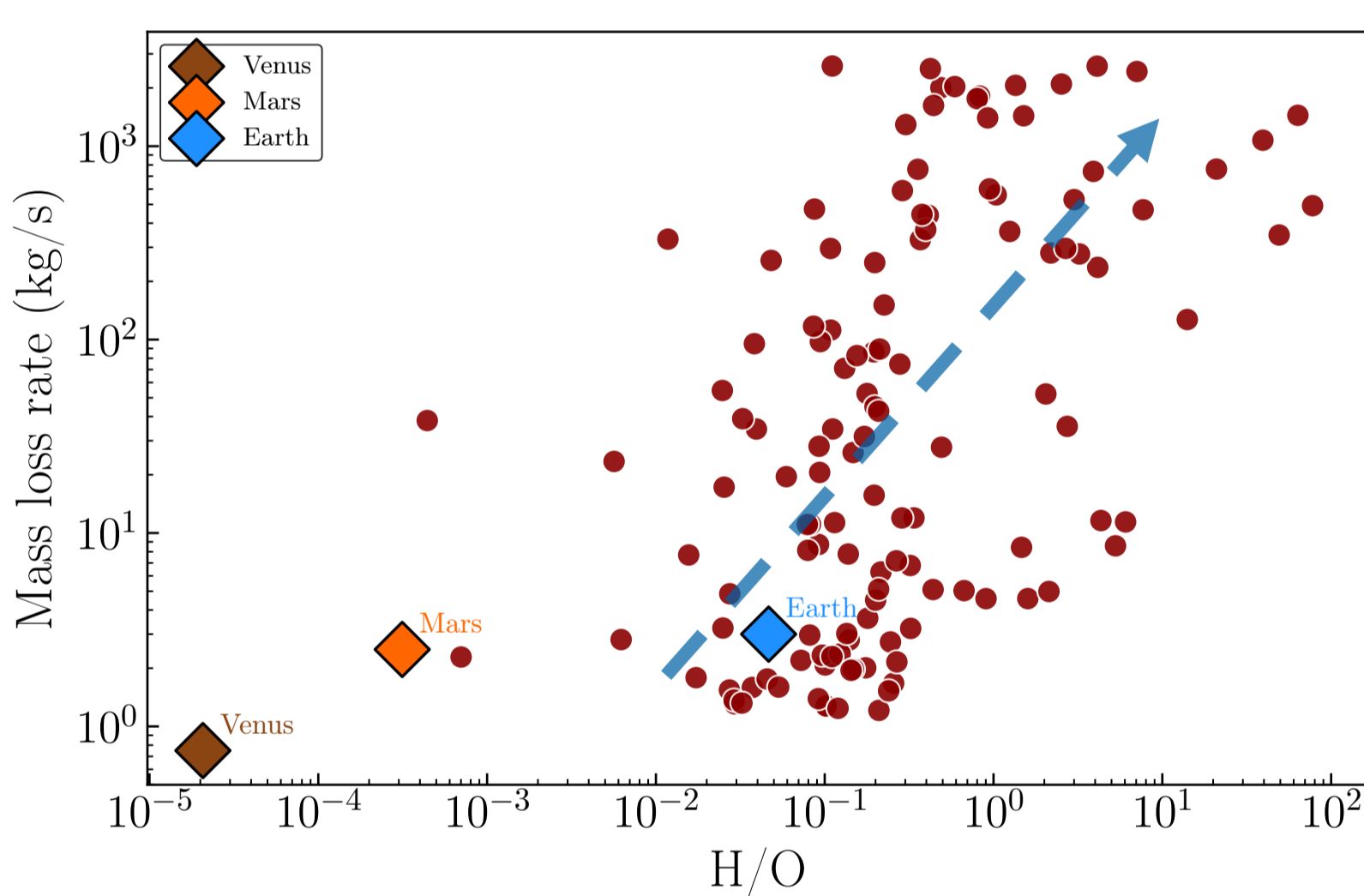
We look at the total amount of mass loss the atmosphere experiences



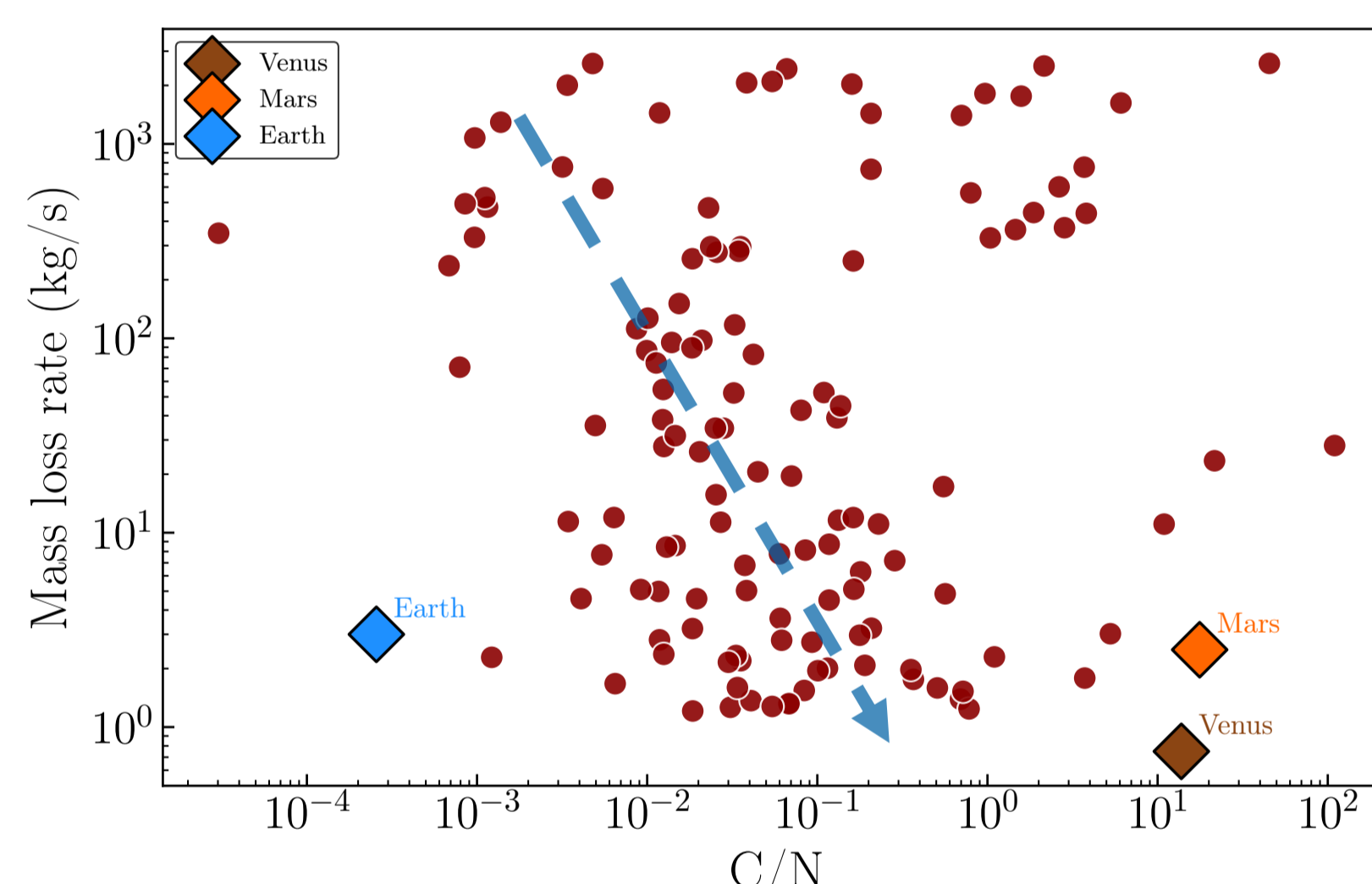
Stable upper atmosphere p-T profile, exobase altitude, final chemical composition

How important are all the elements?

Mass loss increases with the H/O ratio...



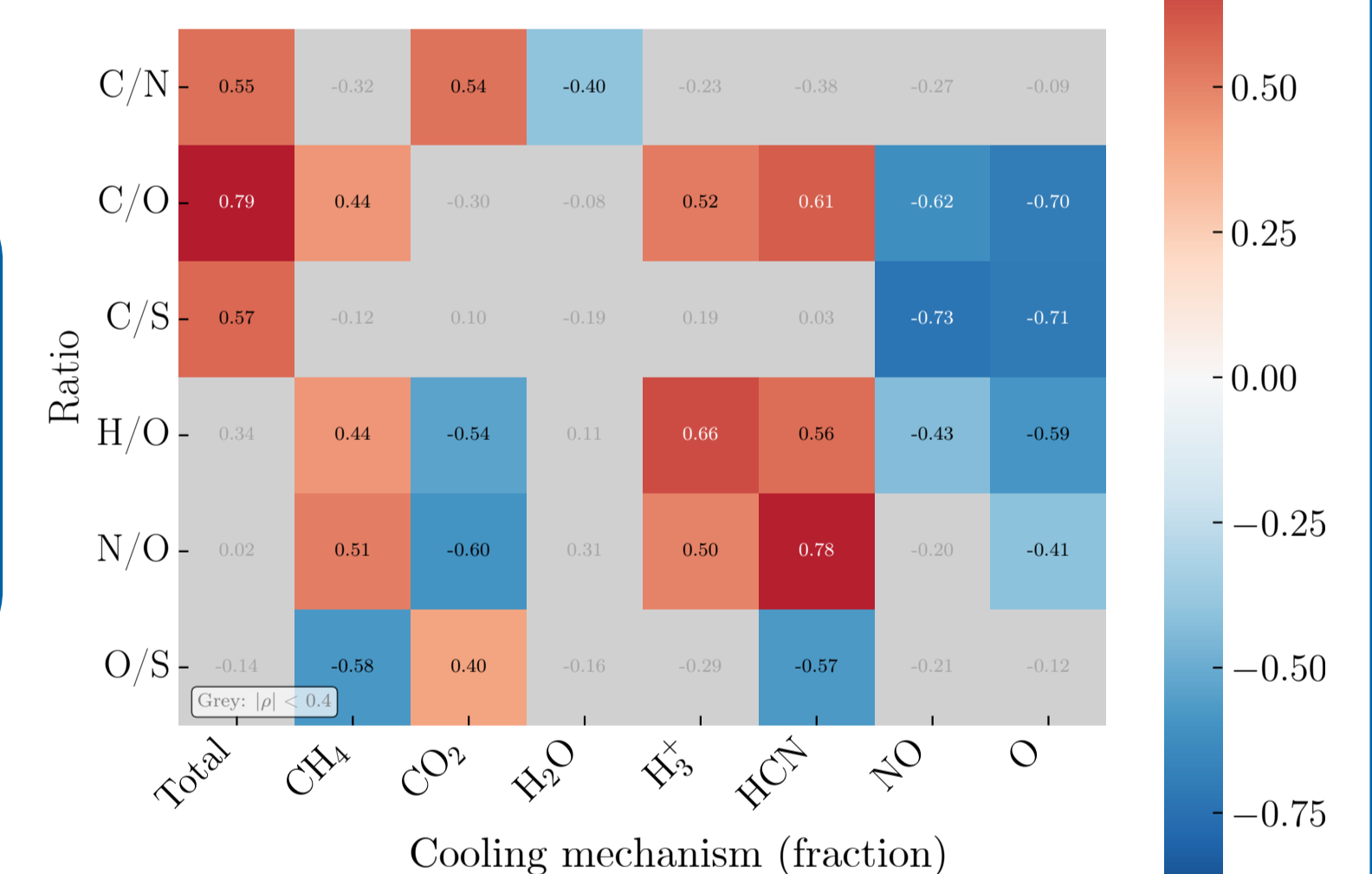
... and decreases with an increase in the C/N ratio



*The high scatter in these figures comes from the inherent chemical interdependence of the species in the model

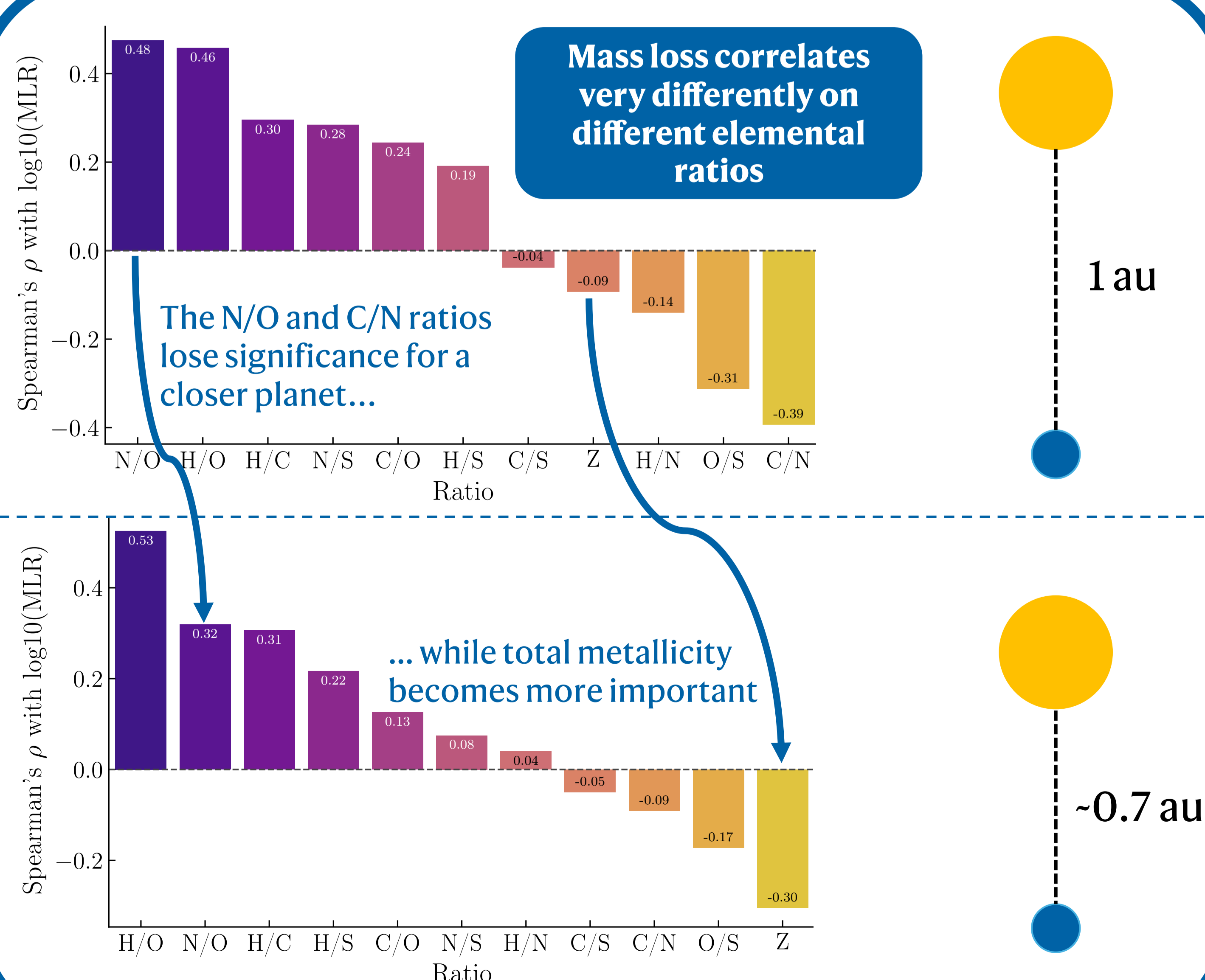
Mass loss is driven by temperature – itself a product of the heating/cooling balance and elemental composition.

Chemical coupling produces surprising correlations. See if you can spot them.*



*for example, the fraction of HCN cooling correlates with the N/O ratio – but it is also anticorrelated with O/S. Similarly H₃⁺ cooling is correlated with C/O, which might be unexpected

The planet-star distance matters too



Takeaways

C/O and metallicity do not give the full picture – we need to consider all other major elemental ratios too.

This will give us:

- better constraints on the thermal structure and potential mass loss
- help us predict which atmospheres we might expect to detect in rocky exoplanets.

References

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